

# Dating a tropical ice core by time frequency analysis of ion concentration depth profiles.

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**Abstract.** Ice core dating is a key parameter for the interpretation of the ice archives. However, the relationship between ice depth and ice age can generally not be easily established and requires to combine a large number of investigations and/or modeling effort. This paper presents a new approach of ice core dating based on time frequency analysis of chemical profiles. We propose a method of counter balancing aliasing that notably improves ice core chronology in comparison with other methods like manual counting. The relationship between the depth of a given ice layer and the date it was deposited is determined using ion concentration depth profiles obtained along a one hundred-meters deep ice core recovered in the summit area of the Nevado Illimani (6350 m asl), located in the Eastern Bolivian Andes (16°37' S, 67°46' W). The results of Fourier time frequency and wavelet transforms are first compared. Both methods are applied to nitrate concentration depth profile. The resulting chronologies are checked by comparison with the multi-proxy year-by-year dating published by de Angelis et al. (2003) and with volcanic tie points. With this first experiment, we demonstrate the efficiency of Fourier time frequency analysis when tracking the nitrate natural variability. Fourier time frequency analysis, then applied to concentration depth profiles of seven other ions, provides information on the suitability of each of these proxies for the dating of tropical Andean ice cores.

## 1 Introduction

Ice core dating, i.e. establishing an univocal relationship between ice depth and age, is not easy, although fundamental for the interpretation of ice archives. Current methods generally combine various approaches like visual stratigraphy, seasonal features allowing to identify annual snow layers, location of specific physico-chemical patterns previously identified in well dated sedimentary or ice

archives, and ice flow or inverse modeling.

Annual layer counting is well suited to the study of high frequency atmospheric processes but requires a certain amount of refinement, due to uncertainty in identifying individual layers. In particular, annual signal may be distorted or smoothed by changes in snow accumulation rate and processes involved in snow deposition and firn aging : snow blowing, wind pumping, water vapor or gradient diffusion and interaction between reactive species may significantly modify the initial registers. Dating uncertainty increases with depth, making it very useful to identify, if any, paleo-events allowing to synchronize different ice cores and potentially to check the current dating for time periods longer than one or two centuries (Narcisi et al., 2005; Traufetter et al., 2004).

The very large number of studies undertaken over polar ice caps since the end of the 1950's has helped to considerably improve our understanding of processes leading to the building of polar ice records and their dating at least for central areas of Antarctica and Greenland. The comparison between deep ice cores recovered in Central Greenland and in Antarctica as well as marine sediment ice cores has provided even more accurate ice chronologies for time periods covering the most recent decades or the past few centuries as well as several hundred of kyr (Ruth et al., 2007; Lemieux et al., 2010). A typical example is the multi-parameter approach combining visual stratigraphy, identification of seasonal maximum along chemical profiles, and age matching with specific climatic and volcanic events used as tie points which provided a year-by-year dating of the GISP2 ice core for the past 100 kyr (Meese et al., 1997). In addition to model time scales, new Greenland time scale has been more recently developed back to 60 kyr before 2000 AD, based on visible stratigraphy and the identification of annual layers along a large set of high resolution records of ice chemical and physical properties (Andersen et al., 2006; Svensson et al., 2005; Rasmussen et al., 2008; Svensson et al., 2008). The thorough comparison of annual layer counting carried out by independent experienced investigators greatly improved the dating accuracy but is a tedious time-consuming effort that highlights the need for automated and less subjective method.

The study of Andean ice records started in the mid 1980s with paleo-climate reconstructions based on the comparison of water stable isotopes profiles with Antarctic or Greenland data (Thompson et al., 1998, 2003). More recently, studies have focused specifically on understanding the atmospheric natural variability at the annual and inter-annual time scale (de Angelis et al., 2003; Correia et al., 2003; Schotterer et al., 2003; Vimeux et al., 2008). However, logistical constraints and harsh weather conditions prevailing over the high altitude cold glaciers of the Andean Cordillera make it often difficult to get the full set of glaciological and meteorological data required by the understanding of processes leading to the building of Andean ice archives. Thus, improving the dating of the cores by developing reliable dating tools remains a key challenge. Particular attention must be paid to sublimation processes and snow remobilization by wind which is suspected to play a prominent role at sites influenced by major cyclonic activity or seasonal convection. Moreover, the rich and intricate fingerprint of Andean volcanism superimposes the long range influence of cataclysmic eruptions

commonly used as time markers along polar ice cores. Nevertheless, most Andean ice cores were  
60 recovered in the tropical belt where the precipitation system is characterized by the yearly alter-  
nation of a warm wet season (around NDJ ) and a colder dry season (around JJA) accompanied  
by significant changes in agricultural practices, soil emissions, and air mass trajectories. Biomass  
burning, increased aridity, and soil dust mobilization send large amounts of aerosol and gases into  
the tropical atmosphere during the dry season at the same time as limited water vapor condensation  
65 and/or riming occur at the surface of falling snow crystals. As a consequence, the concentration  
of chemical species in fresh snow should be significantly higher during the dry season than during  
the wet season. The increasing effect of dry deposition or snow sublimation on the concentration  
of irreversible species in aging surface snow are also likely to be more pronounced during the dry  
season which justifies taking into account seasonal patterns as potential dating tools along ice cores  
70 extracted from tropical Andean cold glaciers. However, seasonal trends in ice cores may be over-  
lapped by other components related to the inter-annual variability of source properties (location,  
emission intensity), of large scale meteorological patterns, and of precipitation occurrence at these  
high altitude sites. Making year-by-year counting independent of the feeling of the observer and  
attempting to minimize the risk of cumulative errors appeared as useful step.

75 Various attempts based on spectral analysis have been developed for automatically exploring sed-  
imentary archives. Preto et al. (2004) for instance used time-frequency analysis to decipher the in-  
fluence of periodic features in Earth current orbit rotational configuration (precession index) on the  
building of sedimentary archives. made of well defined lithofacies unambiguously identified by field  
observation and additional laboratory studies. Smith et al. (2009) present an automated trace element  
80 laminae peak counting procedure based on spectral and wavelet analysis and providing reliable and  
accurate dating of speleothems with rather regular grow rates. Wheatley et al. (2012), developed a  
dating method using the autocorrelation function for large core subsection taken at different depths  
and where missing values or unclear annual signals are reconstructed assuming continuous annual  
layer thickness. An annual layer detection routine based on the framework of hidden Markov models  
85 has been applied by Winstrup et al. (2012) to a stable water isotope record and visual stratigraphy  
of two Greenland cores showing well preserved annual signals at great depth and where manual  
counting has already been carried out. Once all parameters included in the layer model have been  
optimally chosen by the algorithm according with the layer signal in the data (assumed to be well  
constrained), the algorithm reproduces pretty well the manually counted timescales.

90 All these studies focuses on archives containing well marked time related sequences. In partic-  
ular, automated approaches for annual layer counting along ice cores have been applied to records  
showing strong and well constrained seasonal variability and gained at sites fully investigated like  
Antarctic Peninsula or Central Greenland. The context of Andean ice cores is quite different: dating  
tools remain very limited and their accuracy is hard to check.

95 This paper presents a new approach to ice core dating based on the time-frequency analysis of

chemical depth profiles. The method was checked along the upper 100 m of a 137 m deep ice core recovered in 1999 close to the summit of the Nevado Illimani in the Eastern Bolivian Andes (6350 m asl, 16°37' S, 67°46' W). The Illimani ice core was chosen to check the validity of our approach because: (1) continuous high resolution profiles of 13 inorganic ions and 3 carboxylic ions including valuable markers of the main atmospheric reservoirs at a regional scale were available from the surface to 100 m deep; (2) preliminary studies of snow accumulation, of freshly fallen snow, and of the evolution in the composition of surface snow during the dry season were carried out in 2001 and 2002, leading to expect a rather clear seasonal pattern and a good preservation of chemical signals in accumulating snow; (3) a year-by-year dating based on visual stratigraphy was available for the upper part of the core, from the surface to 49.6 m deep, i.e. from 1998 back to 1918 AD (de Angelis et al., 2003). This dating made it necessary to combine the seasonal signals of three wind-borne dust markers (mass of insoluble particles, calcium, and aluminium, (Correia et al., 2003)) and two nitrogen ionic species (nitrate and ammonium). Two pure sulfuric acid peaks were observed at 9.13 m and 33.43 m. They were attributed to the stratospheric fallout in early 1993 and 1964, respectively, of the Pinatubo (June 1991) and Mont Agung (March 1963) eruptions. Debris corresponding to the arrival of tropospheric plumes during the 1991 and 1963 dry seasons were found at 10.8 and 33.4 – 33.8 m, respectively, and used as check points. Dating uncertainty was estimated to be in the range of  $\pm 1$  yr down to 41m (1947 AD) and  $\pm 2$  yr at the bottom of the core section (1919 AD).

Data were processed in two ways: in a first step, the time-depth relationship was deduced from the time frequency analysis of the nitrate profile based on Fourier spectrogram and implemented to construct the ice chronology. Then a similar approach was done using wavelet transforms. We first present the raw data and the method we used to compensate for signal bias induced by the experimental sampling rate which was too low below 68 m and we propose and discuss a preliminary dating. The chronologies resulting from the two spectral methods are compared with the chronology proposed for the upper part of the core by de Angelis et al. (2003) and the volcanic tie points defined by these authors. This provides an assessment of the reliability of the two spectral approaches. In the second step, the Fourier conjugate transform is applied to the concentration depth profiles of seven other ions providing seven other independent ice chronologies. The comparison of these 8 independent dating results between each other and with the tie points used by de Angelis et al. (2003) is discussed in terms of method and marker suitability.

## 2 Raw data

In a first step, the method is described and applied to the nitrate concentration depth profile along the Illimani core,  $c_{NO_3^-}(n)$ . Nitrate is one of the major ions detected in the core and is less sensitive to post-deposition effects than other biomass markers like formate. Moreover, it is expected to

have a strong annual component and to be not markedly influenced by sporadic events like volcanic eruptions or dust storm which can disrupt the seasonal pattern of other ions. Therefore it was considered as most suitable for checking our method. Raw data of  $c_{NO_3^-}(n)$  are presented in Figure 1. To counterbalance firm densification, all depths and layer thickness are expressed in meters of water equivalent ( $m we$ ). The depth  $d(m we)$  expressed in meters of water equivalents and corresponding to a given depth  $d(m)$ , measured along the core and expressed in meters, is given by:

$$d(m we) = \sum_{i=0}^{d(m)} \Delta d(i) * \delta(i) \quad (1)$$

where  $\delta(i)$  is the density of the firm depth increment  $\Delta d(i)$ .

In the left part of Figure 1, we have reported nitrate concentrations as a function of depth along the whole ice sequence. The three panels in the left part of the figure correspond to selected zoomed sections at 25 – 30  $m we$ , 42 – 47  $m we$ , and 62 – 67  $m we$  from top to bottom, respectively. The nitrate depth profile clearly shows an alternation of relative maxima (dry seasons) and minima (wet seasons) corresponding to its annual cycle. The spatial frequency<sup>1</sup> of this cycle increases from the surface to 48  $m we$ .

The apparent decreasing trend with depth of the amplitude and frequency of concentration peaks below 48  $m we$  is due to a change in sampling resolution resulting in inadequate sample length. This under-sampling problem and the approach to address it are discussed in section 4.2.5.

### 3 Method

#### 3.1 Principle

As stated above, our investigation is based on the hypothesis that concentrations of chemical species entrapped in falling snow are periodic in time and that the period is known. We assume that the variability of the chemical composition of tropical Andean precipitation is controlled by strong seasonal changes in atmospheric properties with a time period of one year. However, and in spite of a good preservation of aging snow, the time dependency of chemical signals cannot be directly inferred from concentration depth profiles due to the inter-annual variability of annual net accumulation rate and to ice thinning through firm densification and ice flow: the annual layer thickness,  $a(d)$ , varies with depth. A relationship can be written between  $a(d)$  and the spatial frequency  $f_s(d)$ :

$$f_s(d) = \frac{1}{a(d)} \quad (2)$$

The evolution of the spatial frequency associated with the variable periodicity as a function of depth is determined using Fourier time frequency analysis (Flandrin, 1993; Cohen, 1994) and orthogonal Morlet wavelet transforms (Goupillaud et al., 1985), applied to the nitrate concentration

<sup>1</sup>The spatial frequency, inverse of the annual thickness, is expressed in  $(m we)^{-1}$

depth profile. The spatial frequency  $f_s(d)$  establishes the relationship linking depth and time delay,  $t(d)$ , where  $t(d)$  is the difference between the drilling date and the deposition date of the layer located at the depth  $d$ .

165 In this way, we get the time-depth relationship along the part of the Illimani ice core where continuous ion concentration profiles are available, i.e. from the surface to 76.4 *mwe* depth. A comparison is made between results derived from the Fourier time frequency analysis, the wavelet transforms and a previously published year-by-year dating approach to evaluate the accuracy of spectral analysis methods applied to chemical species. Independent chronologies are then established using nitrate,  
170 formate, oxalate (indicators related to continental biomass), calcium, magnesium, potassium, and sodium (erodible soils) and sulfate (more influenced by volcanism).

### 3.2 Determination of the time-depth relationship

Let  $a(d)$  be the layer thickness corresponding to the depth increase, during a period of one year at the depth  $d$ . To this spatial period, we relate the spatial frequency:

$$175 \quad f_s(d) = \frac{1}{a(d)} = \frac{dt(d)}{dd} \quad (3)$$

Let  $d_{min}$  be the depth origin of the determination of the spatial frequency and  $t(d)$  the time delay from the drilling date to the depth deposition time.

The time delay from drilling date to the deposit time of layer at depth  $d$ :

$$t(d) = \int_{d_{min}}^d f_s(u) du + t(d_{min}) \quad (4)$$

180 gives the relation between time and depth from the spatial frequency. This relation is the central point of our method: it shows that the time depth relationship can be determined from the measurement of the spatial frequency as a function of depth.

## 4 Data processing

As stated above, data processing is first applied to the nitrate concentration depth profile.

### 185 4.1 Data pre-processing

We have a set of 1697 samples of  $c_{NO_3^-}(n)$ . The nitrate concentration function  $c_{NO_3^-}(n)$  is given for a set of depths  $d$  not regularly distributed from 0 to 76.4 *mwe*. This concentration depth profile has high amplitude spikes related to particularly intense biomass burning events during the dry season and its mean value is different from zero. We first threshold these spikes at the mean value of  
190 concentration relative maxima calculated along the profile.

Time frequency analysis requires samples regularly distributed with depth, without high impulses and with a mean value equal to zero. In order to get a concentration profile suitable for time frequency analysis the data set is oversampled by a spline interpolation on 2048 samples regularly distributed in depth. As shown in Figure 2 where  $c_{NO_3^-}(n)$  raw data are reported along with spline interpolation for a 8 *m we* long ice section. The interpolation fits pritty well the original data. Then, the highest pulses are removed by thresholding. The threshold is determined using the cumulative distribution function of measured concentrations. Optimizing the thresholding effect on measured concentrations led us to fit the threshold to the value corresponding to the 0.98 level of the cumulative distribution function. Finally, we remove the mean concentration value (Wheatley et al., 2012).

After spline re-sampling, thresholding, and mean value removal, we estimate the averaged concentration spectrum (Figure 3). The low frequency high level visible in Figure 3 is due to the red character of geophysical noise. These dominant low frequencies would perturb the tracking of the spatial frequency by time frequency analysis. They are eliminated by applying a high-pass filter of cut-off frequency  $f_b$ . The spatial low frequency estimate is sensitive to the value of this cut-off frequency. Section 3.2.4 explains how the cut-off frequency is optimized using volcanic markers<sup>2</sup>.

The normalized new concentration obtained after this pre-processing is called  $S_X(l)$ , where  $X$  is the component under investigation and  $l$  is the sample index ( $l$  varying from 1 to 2048).  $S_{NO_3^-}(l)$  is presented in Figure 4.

The sampling period  $P_s$  resulting from the spline interpolation is:

$$P_s = \frac{76.4}{2048} = 0.0373 \text{ m we} \quad (5)$$

This sampling period enables us to represent the concentration inside the spatial frequency band from 0 to the "Shannon spatial frequency",  $f_{Shannon}$ :

$$f_{Shannon} = \frac{1}{2P_s} = 13.40 \text{ (m we)}^{-1} \quad (6)$$

## 4.2 Estimate of the spatial frequency

The spatial frequency determined here using Fourier time frequency analysis and wavelet transforms, give the energy distribution of concentration profiles in the depth-spatial frequency plane. The time frequency representation based on Fourier transforms leads to spectrograms. Wavelet transforms provide scalograms (Daubechies, 1990; Rioul and Duhamel, 1992). In the next section, we present spatial frequency estimates deduced first from the spectrogram and then from the scalogram.

### 4.2.1 Spectrogram and scalogram

The spectrogram is obtained by pre-windowing the signal around a given depth, calculating its Fourier transform, taking the square modulus, and doing the same for every depth. The spectro-

<sup>2</sup>Alternatively a whitening of the concentration profiles spectra could be applied. We have tried this whitening and we have see that it tends to hidden the one year period trend present in the signal

gram is given by:

$$SPECT_X(d, f_s) = \left| \sum_{l=0}^{N-1} S_X(l)h(l-d)e^{-2j\pi\frac{f_s l}{N}} \right|^2 \quad (7)$$

225 where  $h(d)$  is a sliding window of length  $N$  (Hanning function). The layer thickness resolution of the spectrogram is determined by the thickness of the selected sliding window  $h(d)$ .

Two orthogonal Morlet wavelets (Goupillaud et al., 1985) with  $S$  successive compression factors (scales) are used to calculate the scalogram. At each depth and at each scale we project the concentration  $S_{NO_3^-}$  on the vector of the  $S$  orthogonal Morlet wavelets. The power at each scale is the  
 230 sum of the square of the projection on the two orthogonal Morlet wavelets. This gives the power of the concentration signal on the depth-scale plane (depth-scale representation). From the relation between the scales and the spatial frequencies we get, by a spline transform of the scalogram in the depth-spatial frequencies plane called  $SCAL_X(d, f_s)$ .

Figure 5 shows the spectrogram (bottom) and the scalogram (top) of nitrate concentration in the  
 235 depth spatial frequency plane. The spectrogram is calculated for  $N$  consecutive samples. On the basis of successive tests, we determine that the best plots of the spectrogram are obtained for a window size of 128 samples corresponding to a layer thickness of 4.775 *m we*. For the scalogram we use  $S = 35$  scales that covers continuously the spatial frequency band of the concentration signal.

The depth trend of energy main value, and thus of spatial frequency, is tracked by the green  
 240 "streak" on the spectrogram and on the scalogram shown in Figure 5. Spatial frequency tends to increase from approximately 1  $m we^{-1}$  to 8  $(m we)^{-1}$  along the upper first 48 *m we*. This increase is consistent with annual layer thinning with depth due to ice flowing. Deeper than 48 *m we* the spatial frequency, tracked by the green "streak", decreases. The sudden trend reversal of the spatial frequency and its abnormal decrease below 48 *m we* (already visible in the raw concentration depth  
 245 profile in Figure 1) can be explained by spectrum replication as we will explain in section 4.2.5.

#### 4.2.2 Estimate of the spatial frequency

Spatial frequency estimates from the spectrogram and the scalogram are given by the spatial frequency  $f_s$  of the spectrogram and scalogram maxima at each depth,  $d$ :

$$250 \hat{f}_{s,spec}(d) = \max_{f_s}(SPECT_X(d, f_s)) \quad (8)$$

$$\hat{f}_{s,scal}(d) = \max_{f_s}(SCAL_X(d, f_s)) \quad (9)$$

where  $SPECT_X(d, f_s)$  and  $SCAL_X(d, f_s)$  are the spectrogram and the scalogram corresponding to the component  $X$  (here nitrate), and are expressed as functions of depth and of spatial frequency.

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Figure 6 shows spatial frequencies estimated from the spectrogram (left) and the scalogram (right). Direct estimated spatial frequencies increase from 0 to 48 *m we* and then decrease due to spectrum

replication. Discontinuities are visible on both panels, however their number is significantly greater on the scalogram estimate than on the spectrogram estimate. These discontinuities come from transients. This is because the concentration signal is made of successive events of short duration and varying amplitude and their occurrence is modulated by a seasonal periodicity. It is well known that the wavelets are very sensitive to transients (Mallat et al., 1992), and, as discussed below, the greater number of discontinuities observed in the scalogram led us to consider, here the spectrogram as a better tool than the scalogram for tracking the spatial frequency in our concentration signal.

#### 265 4.2.3 Discontinuity correction with a median filter

Median filters are used to eliminate the discontinuities in spatial frequency estimates,  $\hat{f}_s(d)$  (Coquerez et Philipp, 1995). We use here a median filter of length  $M = 2m + 1$ .

The output of median filters applied to the spatial frequency estimates,  $\hat{f}_s(d)$ , provided by the spectrogram and the scalogram is given by:

$$270 \quad MF[\hat{f}_s(d)] = \text{median}\{\hat{f}_s(i), i \in [d - m, d + m]\} \quad (10)$$

The corresponding results are shown in Figure 6 (lower panels).

#### 4.2.4 Filter optimization

In the pre-processing step (see section 4-1), we have introduced a high-pass filter with a cut-off frequency  $f_b$ . Here we use a median filter of length  $M$ . To determine the optimum values of the high pass filter cut-off frequency,  $f_b$ , and the median filter length,  $M$ , we use two volcanic references: the depths where the volcanic debris attributed to tropospheric deposits from the Pinatubo and the Mount Agung eruptions were found (de Angelis et al., 2003). The Tambora eruption is not taken into account here because the corresponding layer is located below the aliasing depth.

Eruption dates were calculated for the following set of  $f_b$  and  $M$  values:

$$280 \quad f_b = (0, 2 + 0, 1p) \quad (11)$$

$$p = [0, 2, \dots, 10] \quad (12)$$

$$M = [17, 33, 65] \quad (13)$$

285 The optimum values of  $f_b$  and  $M$  are the ones that minimize the distance between estimated dates and historical eruption dates using the criterion:

$$\hat{F}_{(f_b, M)} = |D_{(Pin)} - \hat{D}_{(Pin)}(f_b, M)| + |D_{(Aggu)} - \hat{D}_{(Aggu)}(f_b, M)| \quad (14)$$

where  $D_{(Pin)}$  and  $D_{(Aggu)}$  are the true eruption dates for Pinatubo and Mount Agung, respectively, and  $\hat{D}_{(Pin)}(f_b, M)$  and  $\hat{D}_{(Aggu)}(f_b, M)$  the estimated dates obtained by combining the parameters  $(f_b, M)$ .

When applied to  $S_{NO_3^-}$  spatial frequencies estimated with the spectrogram and the scalogram, the optimization procedure leads to the following optimum values:

Spectrogram estimate:

$$295 \quad f_b = 1.1 \text{ mwe}^{-1} \text{ and } M = 65 \quad (15)$$

Scalogram estimate:

$$f_b = 0.2 \text{ mwe}^{-1} \text{ and } M = 65 \quad (16)$$

Figure 6 shows the spatial frequency estimated with the spectrogram (left) and the scalogram (right) as a function of depth and obtained with the optimum values of  $f_b$  and  $M$ . Discontinuities are smoothed on the two plots. Their attenuation is more efficient on the spectrogram estimate than  
300 on the scalogram estimate.

#### 4.2.5 Evidence of spectrum replication

The initial core sampling was continuous: the core was cut into 1697 successive ice sections and analyses were performed on every ice section once melted. This analytical procedure is equivalent  
305 to an averaging sampler. The thickness, of every ice slice  $d_{slices}$ , which determines the depth period of the sampling is an experimental data varying with depth. At this experimental depth period, we associate a "chemical Shannon frequency"  $f_{chem}$ :

$$f_{chem} = \frac{1}{2 * d_{(slices)}} \quad (17)$$

We present in Figure7 the "chemical Shannon frequency" (dark line),corresponding to the true  
310 core sampling, the "spline Shannon frequency" (dark dotted line) corresponding to the regular resampling applied to original data (see section 4.1) and the spatial frequency estimate obtained previously and corresponding to the one year periodicity expected for the data (called "one year frequency": green line).

Three distinct parts can be distinguished in Figure 7. From 0 to approximately 35  $mwe$  (Part 1)  
315 the "effective" Shannon frequency which is the lowest of the two Shannon frequencies reported in Figure 7, is the chemical one and the "one year frequency" is not aliased (Part 1). From 35 to 50  $mwe$  (Part 2) the "effective" Shannon frequency is the "spline Shannon frequency". In this region the "one year frequency" remains not aliased. At 50  $mwe$  we observe that the effective Shannon frequency is again the "chemical Shannon frequency" and that the increasing "one year frequency"  
320 becomes equal to the "chemical Shannon frequency" and, then decreases (Part 3).

This increase of the "one year frequency" results from annual layer thinning with depth in relation with ice flowing. Given that the influence of this thinning effect increases with depth and is certainly stronger below 50  $mwe$  than above, we can conclude that the decreasing trend of the one year frequency observed below 50  $mwe$  is an artefact due to spectrum aliasing.

## 325 4.2.6 Correcting the aliasing: de-aliasing

In order to recover the non-aliased value of the one year frequency  $f_s(d)$  as a function of depth  $d$ , we replace the observed (aliased) one year frequency  $f_a(d)$  below the depth where aliasing occurs by:

$$f_s(d) = 2 * f_C(d) - f_a(d) \quad (18)$$

330 where  $f_s(d)$  is the non aliased one year frequency,  $f_C(d)$  the "Chemical Shannon frequency" and  $f_a(d)$  the "observed aliased frequency".

In order to realize the de-aliasing we have to determine as accurately as possible the depth where the aliasing occurs. This depth, called "crossing depth"  $d_{crossing}$ , is theoretically the depth where estimated spatial frequency and chemical Shannon frequency cross. Given the discontinuous nature  
335 of the data, we determine the crossing-depth as follows:

$$d_{crossing} = Argmin|f_C(d) - f_s(d)| \quad (19)$$

Spatial frequency estimates from spectrogram and scalogram, after de-aliasing are presented in Figure 8 (left and right panel, respectively).

## 5 Chronology deduced from nitrate ( $NO_3^-$ ) concentration depth profile and accuracy analysis

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### 5.1 Chronology deduced from nitrate ( $NO_3^-$ )

From equation (4), the estimate of the time delay between the depth origin of the determination of the spatial frequency,  $d_{min}$ , and the date corresponding to the depth  $d$  is:

$$\hat{t}(d) = \int_{d_{min}}^d \hat{f}_s(u) du + \hat{t}(d_{min}) \quad (20)$$

345 – where  $d_{min}$  coming from the windows of the spectrogram (N samples) and the median filter (M samples) is:

$$d_{min} = \left(\frac{N+M}{2}\right)P_s = 3.6 \text{ m we} \quad (21)$$

– and  $\hat{t}(d_{min})$ , is:

$$\hat{t}(d_{min}) = \hat{f}_s(d_{min}) \cdot d_{min} \quad (22)$$

350 The dates corresponding to the depth  $d$  is:

$$\widehat{date}(d) = d_{drill} - \hat{t}(d) \quad (23)$$

The chronologies respectively derived from the nitrate spectrogram and scalogram are shown in Figure 9. The drilling date  $d_{drill}$ , was fixed at 1999.39 AD. Both chronologies are globally coherent but local discrepancies are visible, particularly near the aliasing edge. This raises the key question of the chronology reliability.

## 5.2 Evaluation of the chronology inferred from $NO_3^-$ concentration depth profile

Evaluating the chronology reliability (accuracy) is a key step to validate our method. For that, we propose three approaches. The first approach refers only to our chronology. In the second approach we use the volcanic markers. Finally, we compare our chronology with the independent year-by-year chronology used by de Angelis et al. (2003), and Correia et al. (2003) for the upper part of the core.

### 5.2.1 Direct evaluation of our method

We have determined the time delay estimates  $\hat{t}(d)$ , with the spectrogram and the scalogram. From the preprocessed nitrate concentration depth profile  $S_{NO_3^-}(d)$  and the time delay estimates  $\hat{t}(d)$  given by the two chronologies respectively resulting from Fourier time frequency and wavelet transforms and using a spline transformation we get the concentration of  $NO_3^-$  as a function of the time delay estimate:  $S_{NO_3^-}(\hat{t})$ .

The spectrograms of  $S_{NO_3^-}(\hat{t})$ , corresponding to Fourier time frequency (spectrogram, top panel) and wavelet transforms (scalogram, bottom panel) are presented in Figure 10 .

These spectrograms show two very different parts: before and after 160 *yr* (before the drilling date, i.e. 1999 AD).

The part of the spectrograms corresponding to time delay values greater than 160 years are associated to the aliased part of the data and, due to the aliasing, all energies are concentrated on frequencies lower than  $1 \text{ yr}^{-1}$ . Thus, we cannot use this part of the spectrograms to check the quality of our method.

For time delay values lower than 160 *yr*, i.e., the non-aliased part of the data, the spectrogram of  $S_{NO_3^-}(\hat{t})$  obtained from the dating with the spectrogram (top panel) clearly shows a track centered at  $1 \text{ yr}^{-1}$ . This indicates that the results of the Fourier time frequency are in agreement with our working hypothesis that aimed to find a seasonal triggering in the nitrate concentration depth profile. On the opposite, an increasing trend from  $0.8 \text{ yr}^{-1}$  to  $1.2 \text{ yr}^{-1}$  is visible between the origin and 160 *yr* on the spectrogram of  $S_{NO_3^-}(\hat{t})$  resulting from the dating with the scalogram (bottom panel). This strongly suggests that wavelet transforms cannot be used to unambiguously detect seasonal variations in our data set.

**Table 1.** Dating of tie points from the spectrogram and the scalogram

Eruption	<i>Pinatubo</i>	<i>MtAgung</i>	<i>Tambora</i>
Historical date	1991.5	1963.3	1815
Date 1 (spectrogram)	1990.9 (-0.6)	1962.4 (-0.87)	1794.1 (-21)
Date 2 (scalogram)	1991.3 (-0.2)	1964.2 (0.94)	1812.5 (-2.5)

### 5.2.2 Evaluation of the chronologies deduced from $NO_3^-$ depth profile by comparison with volcanic tie points

385 The historical dates of Pinatubo, Mt Agung and Tambora tropospheric plume arrivals and corre-  
sponding dates deduced from the use of the spectrogram (Date 1) and the scalogram (Date 2) are  
given in Table 1. As a general rule, specific attention must be paid to the use of volcanic eruptions  
as tie points for the dating of Andean ice cores. Major volcanic eruptions can be clearly identified  
in Antarctic ice cores through the arrival of large amounts of sulfuric acid from the stratosphere.  
390 Nevertheless, the formal identification of a given individual event requires specific micro-physical  
and micro-chemical investigations of ashes to be confirmed unless detailed matching with another  
well dated ice core can be provided. In tropical South America, the quiescent or eruptive activity of  
Andean volcanoes competes with the influence of eruptions taking place farther north in the tropi-  
cal belt. Combined with the lack of information on dispersion and removal parameters as well as  
395 on the relative importance of tropospheric transport and stratosphere-troposphere exchanges, this  
makes uncertain the use of general volcanic markers alone to relate a given peak to a given eruption.  
Nevertheless it must be noted that, the depth corresponding to the plume arrival attributed to Mount  
Agung eruption is in agreement with  $^{37}Cs$  and tritium peaks at 33.5 m (21.7 m we) associated with  
stratospheric fallout in 1964 of major nuclear bomb tests performed in the northern hemisphere.  
400 The dates obtained for the Pinatubo and Mt Agung eruptions with the two approaches used in  
this study are very close to the eruption historical dates. We keep in mind that these reference  
horizons have been used to optimize the high pass filter cut-off frequency and the median filter  
length so it is normal to recover approximately these historical dates. However the results presented  
in Table 1, provide a valuable estimate of the overall precision of every dating method. On the other  
405 hand, the dates deduced from the spectrogram and the scalogram for the arrival of volcanic debris  
attributed to the Tambora eruption by de Angelis et al. (2003) are different from the historical date,  
i.e. 1815<sup>3</sup>. These differences may be due to the fact that this event is located in the aliased part  
of the data, but they must also be considered keeping in mind chronology discrepancies between  
previous studies related to the choice of time references considered to be "certain". The chronology

<sup>3</sup>The closeness of the Tambora datation by scalogram comes from an artefact introduced by the estimation of spatial frequency between 40 and 45 m we (Figure 8, (right panel)).

410 of a nearby ice core based on the counting of pseudo-seasonal layers and the location of major peaks in the electrical conductivity profile (Knüsel et al., 2003) locates at 99.9 *m we* the arrival of a major unknown eruption dated in 1258 AD in Antarctica. The choice of this volcanic anchor layer appeared to be congruent with visual counting. However this chronology was largely reassessed by Kellerhals et al. (2010). Considering thallium as an unambiguous volcanic marker, these authors  
415 used the location of major thallium peaks combined with ice modelling and preliminary  $^{14}\text{C}$  data to locate the 1258 AD horizon at 90.7 *m we*, i.e. 9.2 *m we* higher, which implies a gap of more than 300 yr between both chronologies and underlines the potential bias induced by subjective factors in visual stratigraphy and the choice of tie points.

### 5.2.3 Comparison with a year-by-year multi-proxy counting approach

420 We compare the time scales respectively deduced from the  $S_{\text{NO}_3^-}$  spectrogram and scalogram for the upper part of the core (0 - 36.4 *m we*) with the chronology inferred from the multi-proxy year-by-year dating approach proposed by de Angelis et al. (2003).

The time scale published in 2003 for the upper part of the core (dating 1) is represented by a black line in the upper part of Figure 11 (top panels) while time scales respectively deduced from  
425 the  $S_{\text{NO}_3^-}$  spectrogram (left) and scalogram (right) are represented by a green line (dating 2). The corresponding differences are presented in the bottom part of the Figure. The date gap corresponding to the spectrogram (left panel) remains close to zero from the signal analysis origin depth to 30 *m we*. It slightly increases deeper down, reaching 1.5 years at 35 *m we*. The chronology deviation corresponding to the scalogram (right panel) is higher, particularly between 5 and 15 *m we* and  
430 reaches 4 years at 10 *m we*.

### 5.2.4 Conclusion on the application of Fourier time frequency and wavelet transforms to the $\text{NO}_3^-$ concentration depth profile.

In the previous sections we presented: 1/ a direct evaluation of our chronology; 2/ the comparison of the dates deduced from spectrogram and scalogram analysis for volcanic layers attributed to well  
435 known cataclysmic eruptions and the true eruption dates; 3/ the comparison of time scales provided by scalogram and spectrogram analysis with a previously published chronology based on a multi-proxy year-by-year counting approach.

This allows us to conclude that our method provides chronologies consistent with available information. However, the seasonal variability of nitrate concentrations is clearly identified only when  
440 using the Fourier time frequency. Moreover, the chronology deduced from the spectrogram shows a better general agreement with the time scale published by de Angelis et al. (2003) than the chronology inferred from the scalogram. As shown in Figure 6, this may be explained by the much greater rate of discontinuities introduced when estimating spatial frequencies from the scalogram. It is not relevant to directly evaluate the quality of our method in the aliased part of the data. However, it

445 must be noted that although the Tambora is inaccurately dated, we were able to detect spectrum repli-  
cation due to the sampling procedure, to reverse the aliasing and to propose a preliminary time scale.  
Although not accurate if compared with what may be expected from high resolution concentration  
depth profiles, this time scale is also more reliable than the year-by-year counting performed on the  
under-sampled part of the core as proposed by Ramirez et al. (2003). Indeed, these authors tracked  
450 seasonal variations down to 66 *m we* but with a mean sample length of 5.8 *cm*, i.e. still greater than  
in this study that should have led to significantly overestimate accumulation rates (see also Figure  
1). Using the sulfate layer attributed to the Tambora eruption as an absolute tie point, they estimated  
the uncertainty of their multi-proxy approach to be  $\pm 20$  years at 66 *m we*. This is in the range of  
the date difference found here but fails to take into account the additional error introduced by the too  
455 low sampling rate which is likely to reach several tens of years at this depth.

In the next section, the spectrogram is applied to concentration depth profiles of seven other ions,  
making it possible to compare and discuss the suitability of each ion as a dating proxy.

## 6 Suitability of chemical markers for Illimani ice core dating

The seven ions studied here and compared with nitrate ( $NO_3^-$ ), are sulfate ( $SO_4^{2-}$ ), sodium ( $Na^+$ ),  
460 calcium ( $Ca^{2+}$ ), magnesium ( $Mg^{2+}$ ), potassium ( $K^+$ ), oxalate (*Ox*), and formate (*For*). They  
were chosen for the reservoirs they are expected to originate from and for their emission seasonality.  
The soluble part of continental input at the Illimani site is mainly regional and consists of gypsum  
( $CaSO_4$ ) and halite ( $NaCl$ ). Strong acidic gases of volcanic origin react with wind-blown minerals  
and volcanic ashes leading to the formation of secondary gypsum, halide salts and lower amounts  
465 of other sodium, magnesium and potassium salts. Halite may also originate from the large salars  
present in Northern Chile and Bolivia, mostly during the dry season. In tropical South America, the  
fire season takes place during the dry season, and fire events recorded along the Illimani ice core  
were identified by simultaneous peaks of ammonium, formate and nitrate and significant amounts  
of soluble potassium (de Angelis et al., 2008). Local emissions of nitrate and ammonium by soils  
470 and formate by vegetation may also be expected more likely during the warm wet season. Oxalate is  
directly emitted through biomass burning, but may also be produced through the oxidation of volatile  
organic compounds (VOC) scarcely investigated in the South American atmosphere.

### 6.1 Optimization

Optimal values of the high pass filter cut-off frequency  $f_b$  and the median filter length  $M$  are de-  
475 termined for the seven new ions using the criterion presented in section (4.2.4) with reference to  
Pinatubo and Mount Agung eruption dates. The results of the optimization procedure are summa-  
rized in Table 2, along with the corresponding minimum value of  $\hat{F}(f_b, M)$  defined in Equation  
(17). We also specify if the  $1 \text{ yr}^{-1}$  frequency was found in the corresponding concentration spectra

**Table 2.** Comparison of eight chemical markers

Concentration	Frequency $f_b$	Filter length $L$	$Min \hat{F}_{(f_b, L)}$	Retrieve 1 year
$NO_3^-$	1.1	65	0.27	yes
$O_x$	0.9	65	0.39	yes
$Ca^{2+}$	1	65	0.47	yes
$Mg^{2+}$	0.7	65	0.66	yes
$K^+$	0.7	65	1.06	yes
<i>For</i>	0.2	17	1.25	no
$Na^+$	0.6	17	4.65	no
$SO_4^{2-}$	0.2	33	6.09	no

expressed as a function of the time delay estimate (Figure 12).

480 The eight ions are sorted according to ascending values of the distance criterion  $\hat{F}(f_b, L)$ . This makes it possible to experimentally estimate the relative quality of chronologies based on each concentration depth profile. Peaks corresponding to annual frequency are clearly visible in nitrate, oxalate, calcium, magnesium and potassium panels in Figure 12, where concentration spectra are expressed as a function of estimate time for the eight ions. This is not true for sulfate, sodium and  
485 formate and is consistent with the sorting presented in Table 2 according to the quality criterion.

## 6.2 Chronology overall comparison

We are aware that the in-depth understanding of the spectral characteristics of signals extracted from ion concentration depth profiles requires further investigations to answer many questions, which is not beyond the scope of this paper. However, a very preliminary interpretation of signal analysis in  
490 terms of time marker reliability may be extracted from Figure 13, where time-depth relationships respectively deduced from  $NO_3^-$ ,  $Ox$ ,  $Ca^{2+}$ ,  $Mg^{2+}$  and  $K^+$  concentration profiles are shown along with the time-depth relationship used by de Angelis et al. (2003). A rather good agreement between the five chronologies proposed here as well as with the chronology used by de de Angelis et al. (2003) is observed down to ca 28 *m we*, i.e. from the late 1950s to the date corresponding to the  
495 signal origin depth. Differences increase with depth, reaching a few years at 36 *m we* for  $K^+$  and  $Mg^{2+}$  but with opposite signs. A more detailed representation of chronology trends may be found in Figure 14 which shows the deviation of our eight independent chronologies from the chronology of de Angelis et al. (2003). Nitrate and calcium provide the best agreement with the chronology published by de Angelis et al. (2003) and the two volcanic tie points. This confirms our hypothesis  
500 that these two ions are produced during the dry season: nitrate production is largely dominated by biomass burning compared with soil emission during the warm season and windblown soil dust is by far the most important source of soluble calcium, the relative contribution of volcanic ash remain-

ing low. The good performance of oxalate, both in terms of seasonality and dating accuracy, is very likely due to enhanced production through biomass burning combined with heterogeneous uptake on mineral dust during the dry season as supported by trend similarities in calcium and oxalate deviation curves. Several mechanisms may contribute to overlap the signal seasonality related to emission and/or transport processes of the other ions. Like calcium, magnesium and the part of potassium not related to biomass burning come mainly from erodible soil dust emission, with a sporadic volcanic contribution. However, the composition and dissolution rate of soil dust and volcanic ashes significantly vary depending on soil and eruption types influencing the signal of the less soluble species (i.e. magnesium, potassium) in a more significant way than the signal of the more soluble ones like calcium. Episodic transport of polluted plumes of local origin through convection during the wet season since the mid 1950s in addition to the timing of stratospheric volcanic arrivals are very likely responsible for the absence of clear  $1 \text{ yr}^{-1}$  frequency in sulfate profile and corresponding ice age overestimate. In a similar way, the occurrence of non periodic peaks related to halite formation and to partial dissolution of ashes in volcanic plumes may become significant compared to seasonal halite emission of sodium and thus explain the bad performance of this ion. Despite a similar origin in relation with biomass emissions the formate signal does not show the  $1 \text{ yr}^{-1}$  frequency clearly visible on the nitrate spectrum: formate may be deposited either as salt and then co-vary with nitrogen species, or as formic acid which is known to be sensitive to post deposition processes including peak removal from strongly acidic layers and relocation on mineral particles.

## 7 Conclusions

Spectral analysis appears as a potential ice core dating tool that does not require the preliminary knowledge of net accumulation rate or ice flow characteristics. It is here applied on a tropical Andean ice core recovered in the summit area of the Nevado Illimani (6350 m asl). Fourier time frequency and orthogonal Morlet wavelet transforms are first compared for tracking spatial frequency in nitrate concentration depth profile. High pass and median filters aimed at correcting spectrogram and scalogram discontinuities were optimized using two volcanic layers attributed to the Pinatubo (june 1991) and Mount Agung (march 1963) eruptions by de Angelis et al. (2003). Several results may be highlighted:

- Both methods give the spatial frequency as a function of depth.
- Spectrum replication is detected on both the scalogram and the spectrogram below  $48 \text{ m we}$  and we demonstrate that replication occurs when spatial frequency becomes greater than the experimental Shannon frequency imposed by the initial chemical sampling procedure. We propose a method of counter balancing aliasing that notably improves the ice core chronology in comparison with other methods like manual counting of pseudo-annual layers. However, it has to be underlined that spectrum replication significantly increases the difficulties encountered here, preventing us from directly

assessing the chronology accuracy along the aliased part of the data.

540 - The chronology provided by the Fourier time frequency analysis for the part of the ice core located above the aliasing depth, i.e. above 48 *m we*, is in close agreement with the chronology based on the multi-proxy year-by-year counting of wet season minima and the Pinatubo and Mont Agung eruptions dates. Wavelets transforms appears to be less efficient for the data investigated in this study. The Fourier time frequency analysis is then applied to the concentration profiles of seven other ions, providing eight independent chronologies and information on the natural variability of  
545 the corresponding ions.

This chronology approach seems particularly suitable for ice cores extracted at poorly documented sites, where the dynamics of seasonal signals may be significantly influenced by source proximity or regional meteorological features, and where emissions of local or regional origin make it difficult to unambiguously identify particular events of global importance or anthropogenic trends generally  
550 used as adjusting marks along polar ice cores. Further improvements of this method should result from its application to other archives in relation with geochemical considerations.

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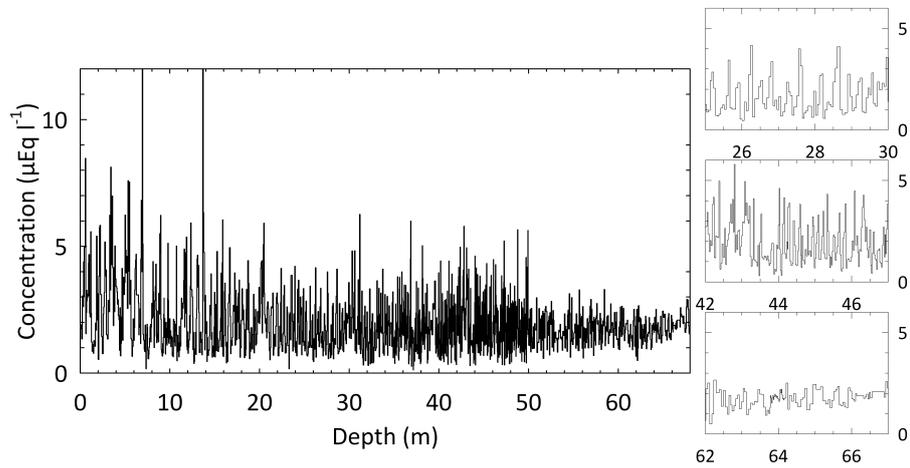
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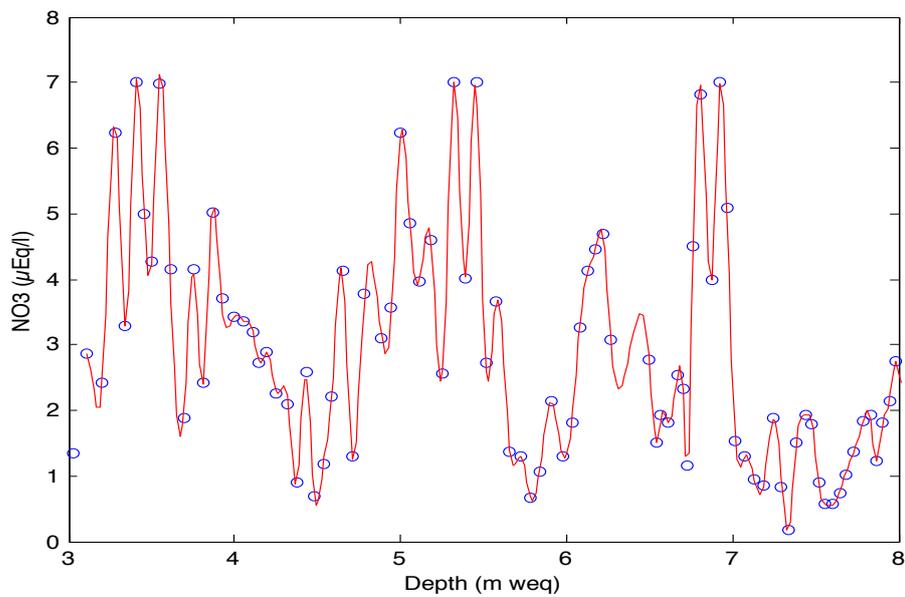
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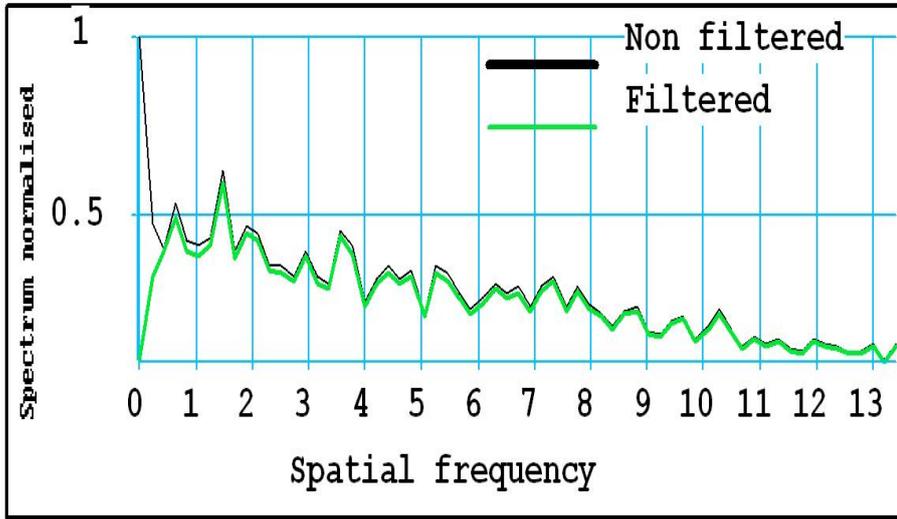
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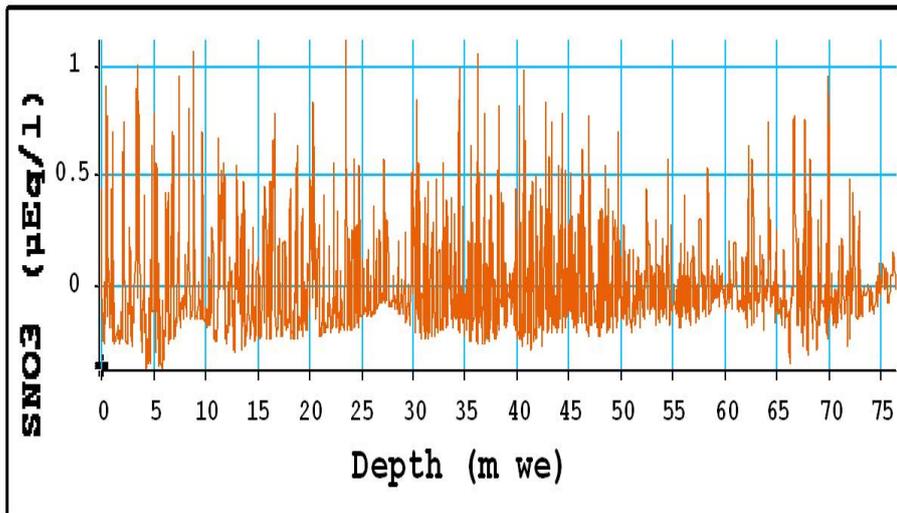
**Fig. 1.** Raw data of  $c_{NO_3^-}(n)$  showing the concentration from 0 to 70 *m we* depth.



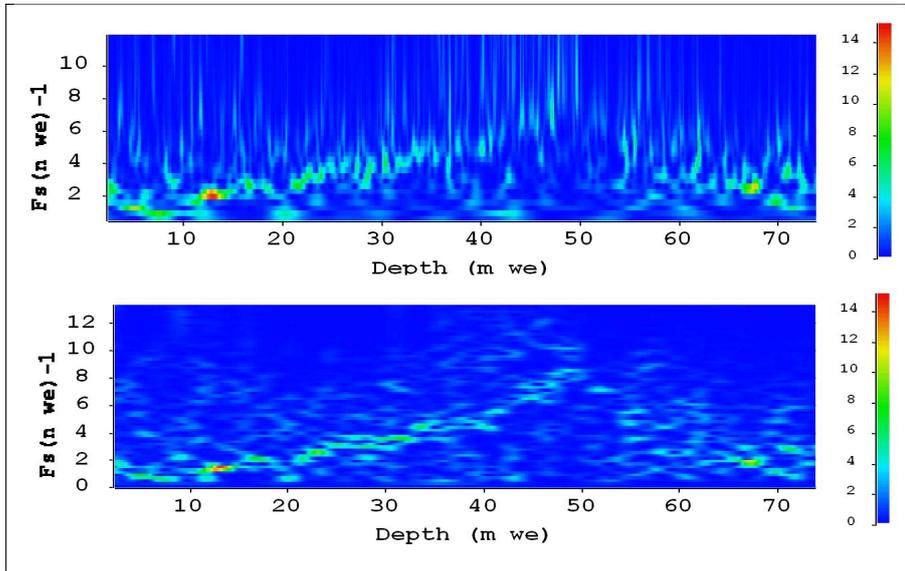
**Fig. 2.**  $c_{NO_3^-}(n)$  raw signal (circles) and the spline interpolation (red line).



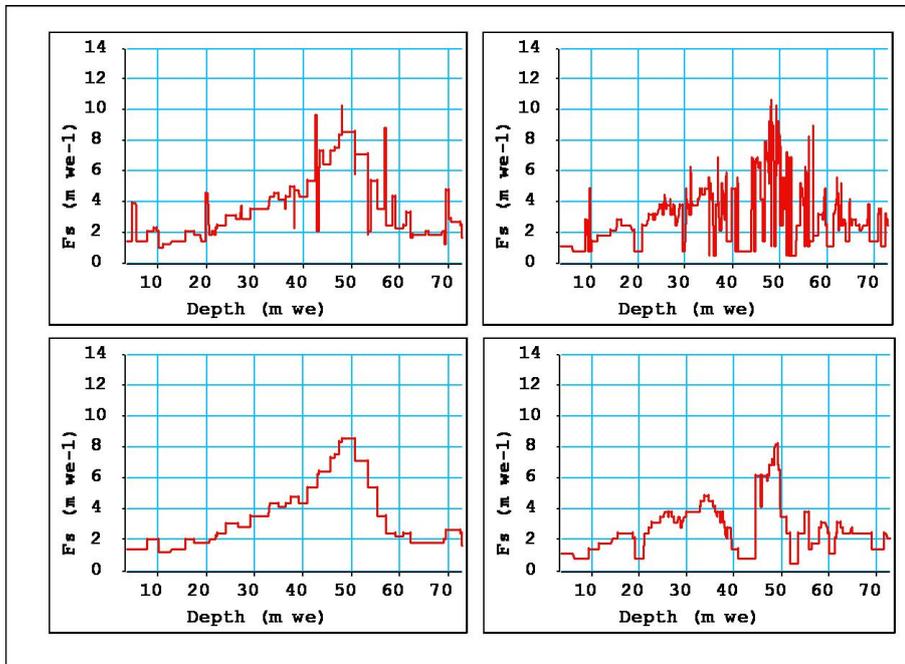
**Fig. 3.** Normalized spectrum as function of the spatial frequency of the  $NO_3^-$  concentration depth profile after resampling and thresholding (black line). The green line gives the spectrum of  $NO_3^-$  concentration depth profile after resampling, thresholding and **high pass filtering**.



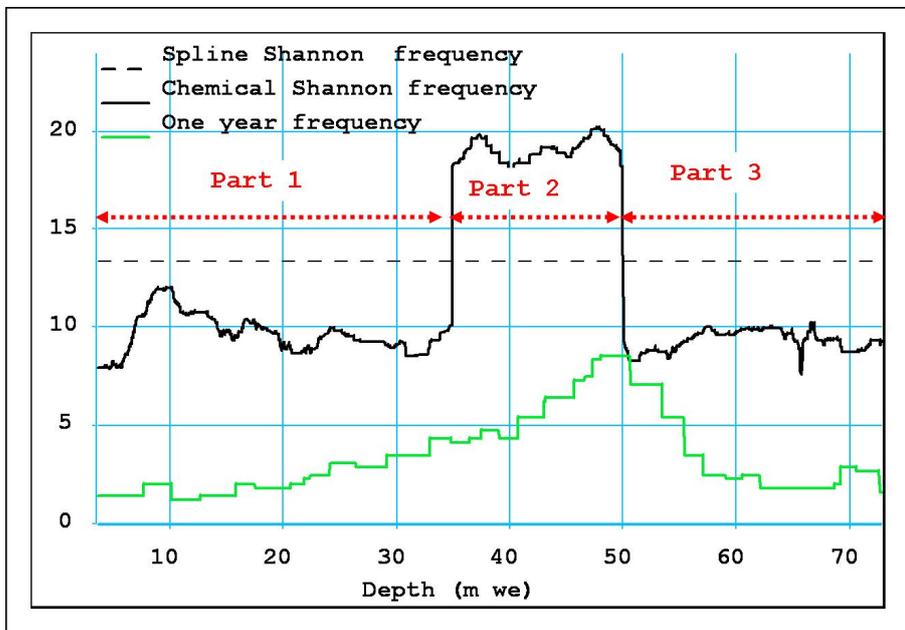
**Fig. 4.** Illimani centred thresholded nitrate concentration  $S_{NO_3^-}$  as a function of depth resampled by a spline interpolation (2048 samples). Concentration is expressed in micro-equivalents per liter,  $\mu\text{Eq/l}$ .



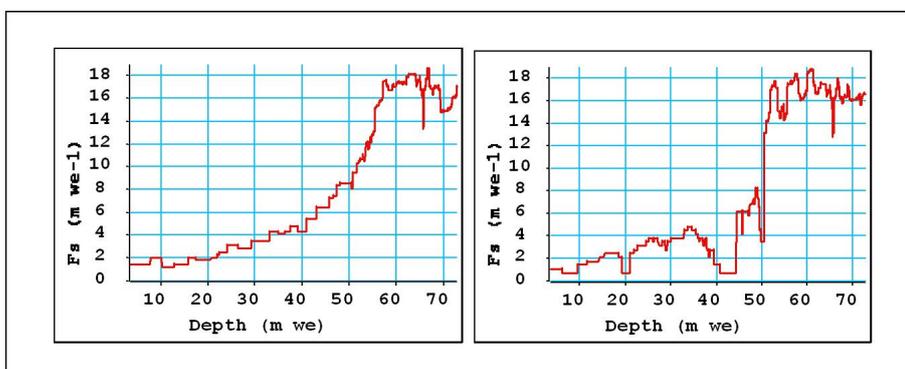
**Fig. 5.** Illimani scalogram of  $S_{NO_3^-}$  (top), Illimani spectrogram of  $S_{NO_3^-}$  (bottom) as a function of depth in  $mwe$  and spatial frequencies  $fs (m we)^{-1}$ . The color scale (from green to red) shows increasing energy of the concentration signal (red pixels correspond to the highest energy value of the figure).



**Fig. 6.** Direct estimates, (up panels) and after application of the median filter (down panels), of  $S_{NO_3^-}$  concentration spatial frequency  $fs (m we)^{-1}$  as a function of depth ( $m we$ ) from spectrogram (left panels) and scalogram (right panels).



**Fig. 7.** Shannon frequency deduced from the true core sampling (“chemical Shannon frequency”, dark line); Shannon frequency deduced from the spline pre-processing (“spline Shannon frequency”, dark dotted line); estimate spatial frequency provided by the spectrogram (“one year frequency”, green line).



**Fig. 8.** Estimates of  $S_{NO_3^-}$  concentration spatial frequency  $f_s(m\ we)^{-1}$  as a function of depth ( $m\ we$ ) from spectrogram (left panel) and scalogram (right panel) after application of the median filter and de-aliasing.

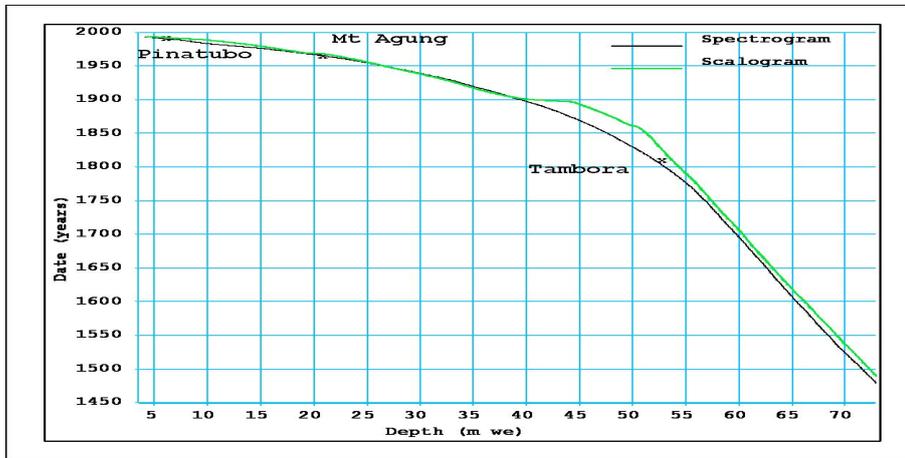


Fig. 9. Illimani chronologies deduced from the  $NO_3^-$  concentration depth profile using the spectrogram (black line) and the scalogram (green line) .

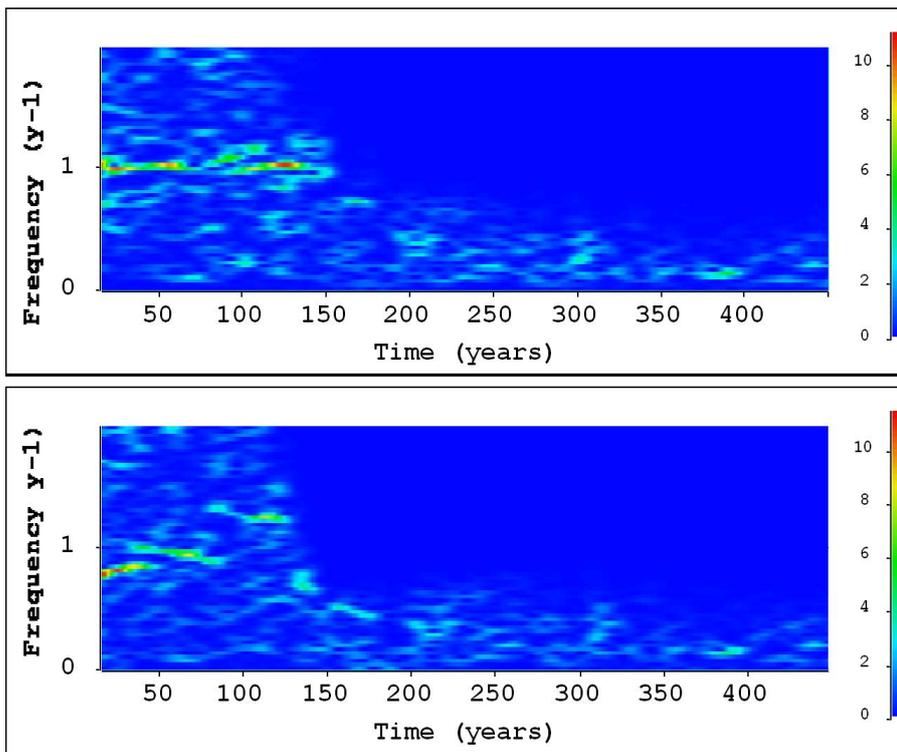
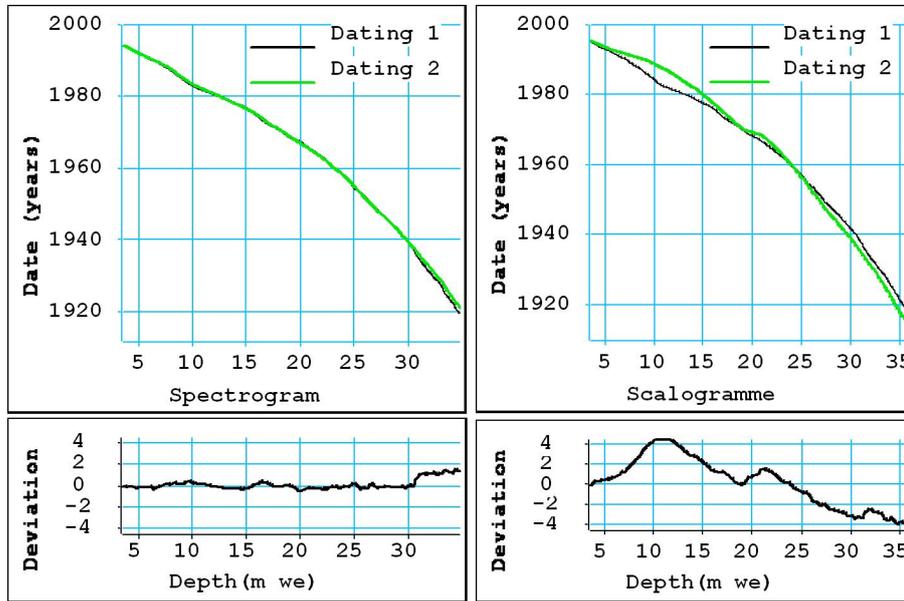
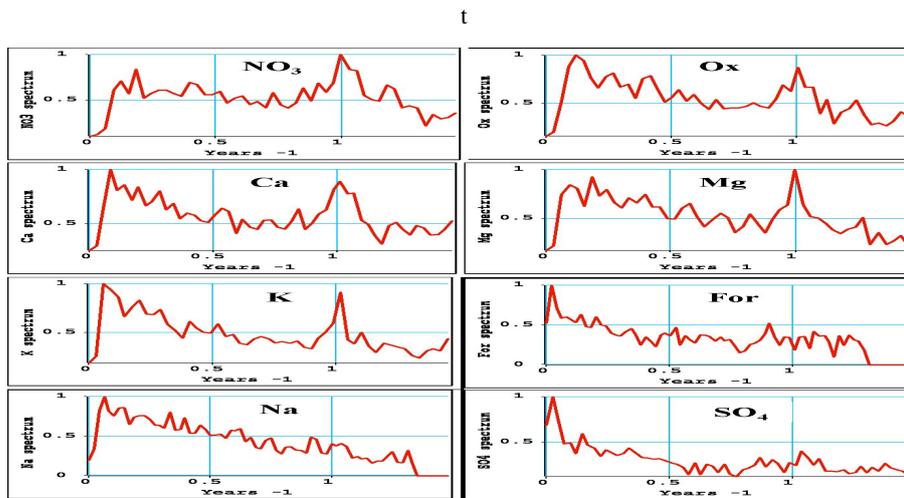


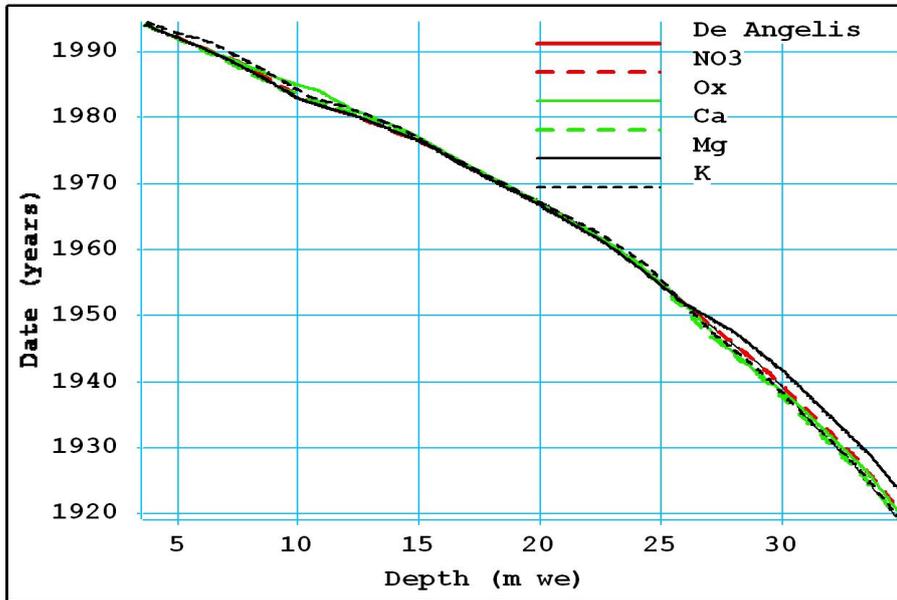
Fig. 10. Spectrograms of  $S_{NO_3^-}$  as a function of time delay estimate  $\hat{t}$ . Top panel: the time delay estimate  $\hat{t}$  is given by the chronology deduced from the spatial frequency estimated with the spectrogram. Bottom panel: the time delay estimate  $\hat{t}$  is given by the chronology deduced from the spatial frequency estimated with the scalogram.



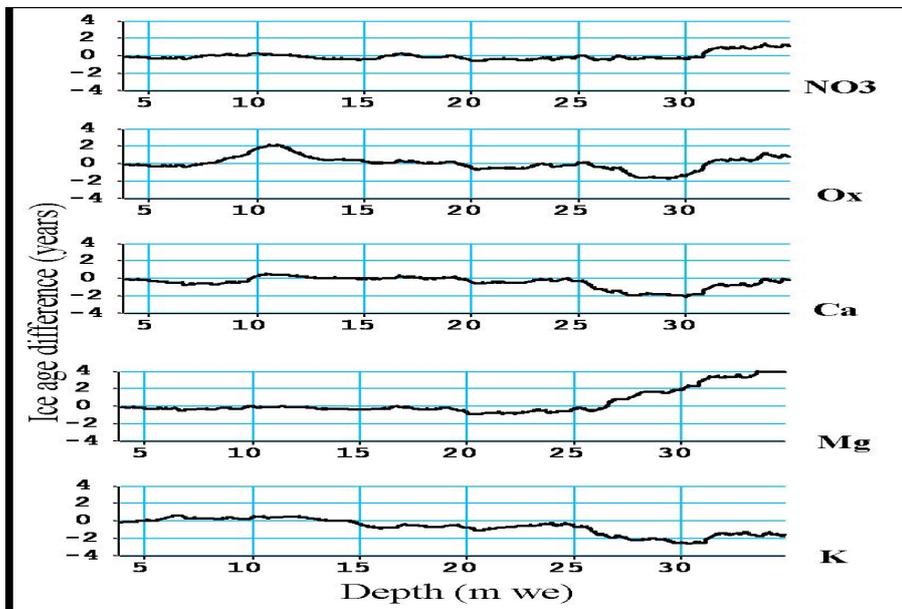
**Fig. 11.** Comparison between the chronology used by Correia et al. (2003) and de Angelis et al. (2003) (dating 1) and chronologies (dating 2) established using the  $S_{NO_3^-}$  spectrogram (left) and the  $S_{NO_3^-}$  scalogramme (right). Dates are expressed as a function of depth from the signal origin depth to 36 *m we*.



**Fig. 12.** Spectra of  $NO_3^-$ ,  $O_x$ ,  $Ca^{2+}$ ,  $Mg^{2+}$ ,  $K^+$ ,  $For$ ,  $Na^+$ , and  $SO_4^{2-}$  concentration depth profiles as a function of time delay estimate.



**Fig. 13.** Comparison of chronologies deduced from  $NO_3^-$ ,  $O_x$ ,  $Ca^{2+}$ ,  $Mg^{2+}$  and  $K^+$  concentration profiles with the chronology used by Correia et al. (2003) and de Angelis et al. (2003). Dates corresponding to ice layer deposits (yr AD) are expressed as a function of depth (*m we*).



**Fig. 14.** Ice age differences (*yr*) calculated for the 36 upper meters (*m we*) between the chronology given by Correia et al. (2003) and de Angelis et al. (2003) and chronologies deduced from  $NO_3^-$ ,  $O_x$ ,  $Ca^{2+}$ ,  $Mg^{2+}$  and  $K^+$  concentration profiles (from top to bottom, respectively).